How can we separate microphysical and dynamical impacts in weather forecast? *A novel modeling methodology Piggybacking* 

> Noémi Sarkadi Pécs, 23. March 2018 Friday Afternoon Cloud Physics Talk

## Outline

- Numerical Weather Forecast basics
- Microphysical and dynamical interactions
- Piggybacking method
- First results in idealized simulations

#### Numerical Weather Forecast - basics

• Hidro-termodynamical equation system (HTES):

$$\frac{du}{dt} = \frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} + w \frac{\partial u}{\partial z} = -\frac{1}{\rho} \frac{\partial p}{\partial x} + fv - lw + F_{sx} \qquad f$$

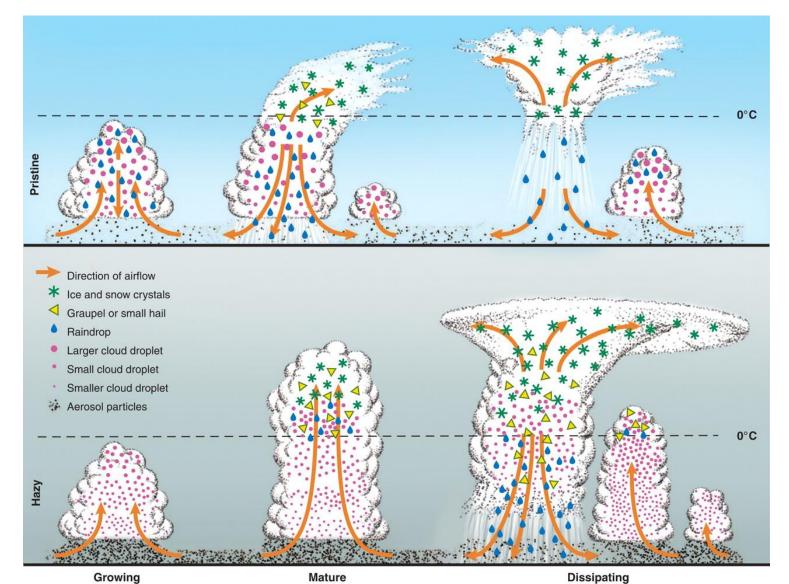
$$\frac{dv}{dt} = \frac{\partial v}{\partial t} + u \frac{\partial v}{\partial x} + v \frac{\partial v}{\partial y} + w \frac{\partial v}{\partial z} = -\frac{1}{\rho} \frac{\partial p}{\partial y} - fu + F_{sy} \qquad \frac{d}{dt}$$

$$\frac{dw}{dt} = \frac{\partial w}{\partial t} + u \frac{\partial w}{\partial x} + v \frac{\partial w}{\partial y} + w \frac{\partial w}{\partial z} = -\frac{1}{\rho} \frac{\partial p}{\partial z} - g + lu + F_{sz} \qquad \frac{d}{dt}$$

$$\frac{d\rho}{dt} = -\rho \nabla V \qquad \frac{\partial \rho}{\partial t} = -\nabla \rho V \qquad \frac{\partial \rho}{\partial t} = -\nabla \rho V \qquad \frac{\partial \rho}{\partial t} = -\nabla \rho V \qquad \frac{\partial \rho}{\partial t} = -\rho \left(\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} + \frac{\partial w}{\partial z}\right) \qquad \frac{\partial \rho}{\partial t} = -\rho \left(\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} + \frac{\partial w}{\partial z}\right)$$

$$\begin{split} f &= 2\Omega\sin\varphi \qquad l = 2\Omega\cos\varphi \\ p\alpha &= RT \\ \frac{d\Theta}{dt} &= \frac{\partial\Theta}{\partial t} + u\frac{\partial\Theta}{\partial x} + v\frac{\partial\Theta}{\partial y} + w\frac{\partial\Theta}{\partial z} = -\frac{\Theta}{T} \bigg( \frac{L_{iv}M_{iv}}{c_{pm}\rho_m} + \frac{L_{iv}M_{iv}}{c_{pm}\rho_m} + \frac{L_{iv}M_{ii}}{c_{pm}\rho_m} \bigg) + \frac{\Theta}{T}D_T + \frac{\Theta}{T}\frac{1}{c_{pm}}Q_R \\ \frac{d\rho_v}{dt} &= \frac{\partial\rho_v}{\partial t} + u\frac{\partial\rho_v}{\partial x} + v\frac{\partial\rho_v}{\partial y} + w\frac{\partial\rho_v}{\partial z} = -\rho_v\operatorname{div}\mathbf{V} + M_{iv} + M_{iv} + F_v + D_v \\ \frac{d\rho_w}{dt} &= \frac{\partial\rho_w}{\partial t} + u\frac{\partial\rho_w}{\partial x} + v\frac{\partial\rho_w}{\partial y} + w\frac{\partial\rho_w}{\partial z} = -\rho_w\operatorname{div}\mathbf{V} - M_{iv} + M_{ii} + S_w + F_w \\ \frac{d\rho_i}{dt} &= \frac{\partial\rho_i}{\partial t} + u\frac{\partial\rho_i}{\partial x} + v\frac{\partial\rho_i}{\partial y} + w\frac{\partial\rho_i}{\partial z} = -\rho_i\operatorname{div}\mathbf{V} - M_{iv} - M_{ii} + S_i + F_i \end{split}$$

### Microphysical and dynamical interactions



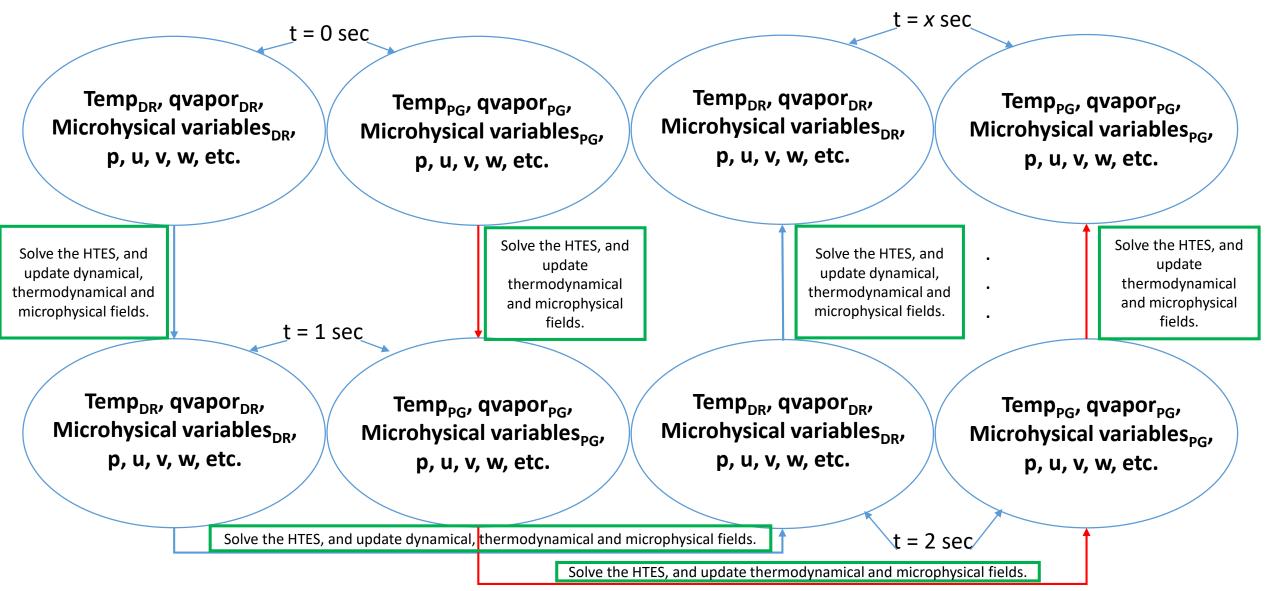
Evolution of deep convective clouds developing in the pristine (top) and polluted (bottom) atmosphere. Cloud droplets coalesce into raindrops that rain out from the pristine clouds. The smaller drops in the polluted air do not precipitate before reaching the supercooled levels, where they freeze onto ice precipitation that falls and melts at lower levels.

The additional release of latent heat of freezing aloft and reabsorbed heat at lower levels by the melting ice implies greater upward heat transport for the same amount of surface precipitation in the more polluted atmosphere.

This means consumption of more instability for the same amount of rainfall. The inevitable result is invigoration of the convective clouds and additional rainfall, despite the slower conversion of cloud droplets to raindrops.

Daniel Rosenfeld, Ulrike Lohmann, Graciela B. Raga, Colin D. O'Dowd, Markku Kulmala, Sandro Fuzzi, Anni Reissell, Meinrat O. Andreae, 2008: Flood or Drought: How Do Aerosols Affect Precipitation?, *Science* 

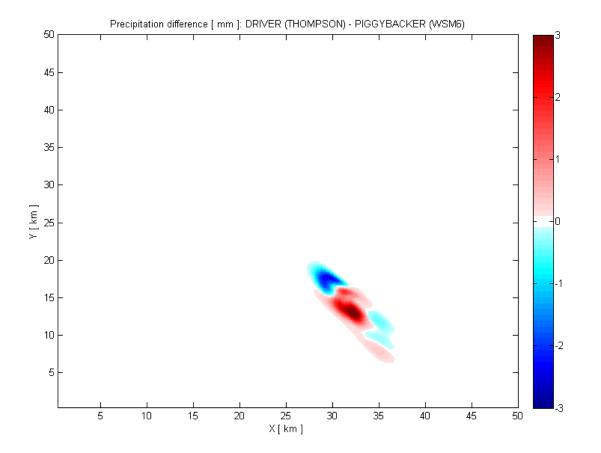
## Piggybacking method



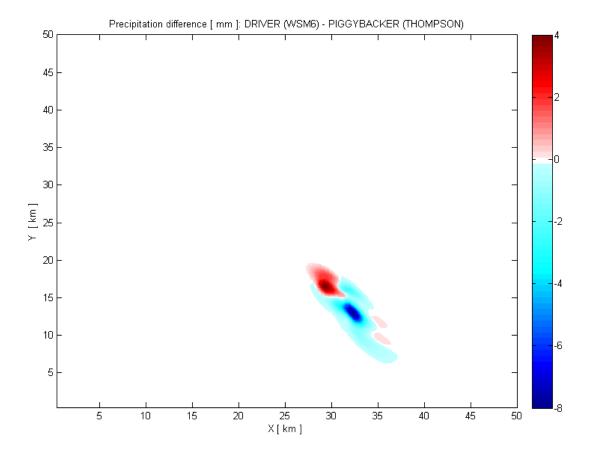
# Results

Idealized simulations with Weather Research and Forecasting (WRF) model

# Piggybacking with different microphysics schemes – Thompson DRIVER, WSM6 PIGGYBACKER

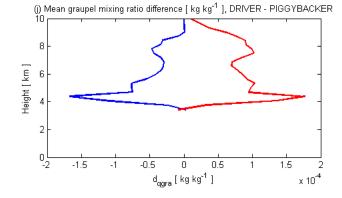


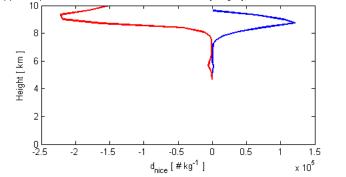
# Piggybacking with different microphysics schemes – WSM6 DRIVER, Thompson PIGGYBACKER

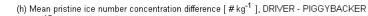


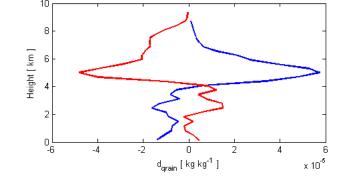
# Difference fields

## Thompson<sub>DR</sub> – WSM6<sub>PG</sub> WSM6<sub>DR</sub> – Thompson<sub>PG</sub>









(g) Mean pristine ice mixing ratio difference [ kg kg<sup>-1</sup> ], DRIVER - PIGGYBACKER

2

3

4

-5

x 10<sup>-5</sup>

Height [ km ] 4

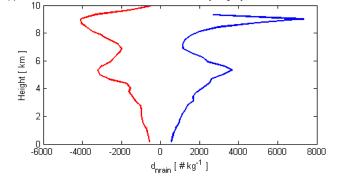
> οL -2

-1

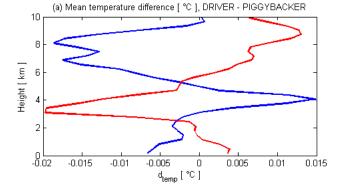
0

1

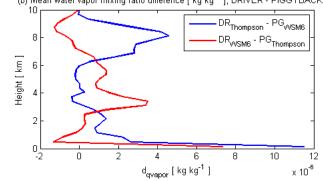
d<sub>qice</sub> [kg kg<sup>-1</sup>]



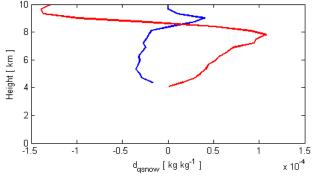




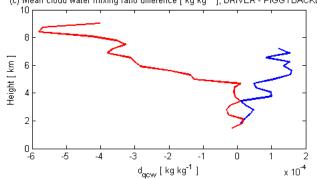
(d) Mean rain water mixing ratio difference [ kg kg<sup>-1</sup> ], DRIVER - PIGGYBACKER



(b) Mean water vapor mixing ratio difference [ kg kg<sup>-1</sup> ], DRIVER - PIGGYBACKER



(f) Mean snowflakes mixing ratio difference [ kg kg<sup>-1</sup> ], DRIVER - PIGGYBACKER



(c) Mean cloud water mixing ratio difference [ kg kg<sup>-1</sup> ], DRIVER - PIGGYBACKER

#### Piggybacking in WRF - Possibilities

- **Option 1:** Using same microphysics as a driver and as the piggybacker, but with different initial conditions (temperature and moisture).
  - Why does it good for us?
  - We can separate the effect of the different micropyhsics on precipitation formation.
  - This is the way how we can test the CCN concentration impacts on cloud formation.
- **Option 2:** Using the same initial conditions, but with two different sets of microphysics
  - Why does it good for us?
  - We can test how the different initial conditions affect the cloud and precipitation formation → investigation of climate change.

### Thank You for Your Attention! Questions?